



Supplementary Materials for

**The Bright Optical Flash and Afterglow from the Gamma-Ray Burst
GRB 130427A**

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Supplementary Text

Modeling

We modeled the GRB 130427A multi-wavelength afterglow emission using the standard external-shock framework (16), where a relativistic forward-shock is driven into the ambient medium by GRB ejecta and that interaction also drives a reverse shock through the ejecta. The 1-dimensional dynamics of each shock is calculated from conservation of mass, energy, and momentum, using the shock jump-conditions (25). The post-shock typical electron energy and magnetic field are calculated assuming that they acquire a certain fraction (micro-physical parameters) of the dissipated energy. Each shock produces synchrotron and inverse-Compton emissions, which are then integrated over the propagation of the shock. The self-absorption and cooling frequencies of the synchrotron spectrum are also calculated from the electron distribution and magnetic field. The best-fit to the afterglow data is then obtained by minimization of chi-square between the sum of model emission components and the observations. For the best fitting models, the effects of synchrotron self-absorption were found to not be important at the observed frequencies during the early epochs. Further, the self-Compton flux from the shocks were also found to be smaller at all times and all frequencies than the corresponding synchrotron fluxes shown in figure 4. However, we did find that the inverse Compton process does play an important role because dominates the electron radiative cooling and determines the synchrotron efficiency above the 100 MeV cooling frequency. The cross-components of synchrotron photons produced by one shock and Compton up-scattered in the other have not been not calculated, but we expect them to be typically smaller than the joint synchrotron/self-Compton flux from the shock.

Our models that successfully fit the multi-wavelength emission require a reverse-shock that is sustained by energy injection throughout the prompt and early afterglow phases. The nature of this injection must vary in three episodes that signal a change in the dynamical and micro-physical parameters of the reverse shock. The best fit parameters for the initial time interval are flash reverse-shock (fRS)-- onset at 4 s, end at 15 s, incoming ejecta Lorentz factor 730, magnetic field parameter 0.008, electron energy parameter 0.006, index of electron power-law distribution with energy 1.9. To account for the transition from the rapid optical flash decay to the more gradual optical afterglow decay observed after 100 seconds requires a change on the reverse-shock dynamics and micro-parameters at ~ 15 seconds. After each change, we calculate the adiabatic and radiative cooling of the ejecta electrons of the previous injection episode. When those electrons cool to an energy below that for radiating synchrotron at 100 MeV, the 100 MeV flux from the 1/Gamma fluid moving toward the observer would display an exponential cutoff in time. However, the fluid moving at larger angles yields a brighter flux after the exponential cutoff, and shows a steep power-law flux decay. The steep power-law decay of the "large angle emission" is seen at 100 MeV after 20 s and after 3 ks in Figure 4. We find the best fit parameters for this new interval are reverse-shock energy injection starting at ~ 15 s and continuing to ~ 3 ksec with a Lorentz factor of 1800, magnetic field parameter of 0.0010, electron energy parameter of 0.012, and electron power-law index of 2.0.

Observations

The three RAPTOR All-Sky Monitors were the first to detect the optical emission from GRB130427A. The burst location was visible at both the Maui and Los Alamos sites at an elevation of 80° and 53° respectively. Each All-Sky Monitor consists of five unfiltered 24 mm f1.5 Canon EF lenses combining to cover about 90% of the sky above 12° elevation. Individually, each All-Sky Monitor obtains a 10 s exposure every 20 s with staggered observing schedules to provide better temporal coverage as a system. We corrected the All-Sky Monitor images using bias and dark frames obtained later that evening as well as flat fields which were created using a median of twilight frames obtained during cloudless evenings during the first half of 2013. The corrected images were then reduced to object lists using the SExtractor package (26) and astrometrically calibrated using ~ 8600 common field stars from the Tycho-2 (27) catalog with SDSS r' band synthetic magnitude estimates (28). Due to sky gradients caused by a nearly full moon about 75° away, these object lists were then photometrically re-calibrated using ~ 55 common Tycho-2 stars within 3° of the burst location.

The narrow field instruments located at the Fenton Hill Observatory in Northern New Mexico, RAPTOR-S and RAPTOR-T, began observing the burst location shortly after receiving the Swift trigger. The RAPTOR-S system consists of a single unfiltered 0.4 m telescope and the RAPTOR-T system consists of four co-aligned 0.4 m telescopes obtaining simultaneous images in four photometric bands (SDSS g', r', i', and z'). Our narrow field observations began at $T-T_0=132.96$ s and the GRB counterpart was clearly visible in all five telescopes. The response sequence of each telescope consists of nine 5 s images followed by twenty 10 s images and 30 s images thereafter. The response images were calibrated using bias and dark frames obtained later that evening as well as flat frames created using a combination of recent twilight images and sky patrol images. The images were reduced to object lists using the SExtractor package and then calibrated using ~ 60 common field stars from the appropriate band in the SDSS DR9 catalog (29). The unfiltered RAPTOR-S data were calibrated to the SDSS r band values.

Our spectral energy distributions (SEDs) in the optical energy range were constructed using the g', r', i', and z' fluxes averaged over four time intervals with end points at 138, 270, 550, 3000, and 7571 seconds after the burst. We applied flux corrections based on extinction curves in Pei (30) to account for $E_{B-V} = 0.0202$ of Galactic reddening (8) and an $A_V = 0.18$ dust column in the host (31). The best fit power law spectrum has a mean slope of $\beta = -0.70 \pm 0.05$ until about 3000 seconds after the burst, when it becomes slightly bluer ($\beta = -0.59 \pm 0.05$).

Table S1.

RAPTOR observations of GRB130427A. T_{mid} is the mid-exposure time in seconds since the GBM trigger (07:47:06.42 UT). The RAPTOR-T data were calibrated using field stars from the SDSS DR9 catalog (29). The RAPTOR All-Sky Monitor data were calibrated against the Tycho-2 Catalog SDSS r' estimates (28)

T_{mid}	t_{exp}	C_r	g'	r'	i'	z'	Telescope
(s)	(s)	(mag)	(mag)	(mag)	(mag)	(mag)	(name)

RQD2—All Sky Monitors							

-9.75	10.00	> 10.1	D03
-4.33	10.00	> 9.9	D02
1.81	10.00	> 9.3	D01
5.86	10.00	8.421+0.075-0.070	D03
14.31	10.00	7.030+0.030-0.029	D02
21.48	10.00	7.501+0.032-0.031	D03
22.21	10.00	7.625+0.061-0.058	D01
33.27	10.00	8.752+0.098-0.090	D02
37.09	10.00	8.525+0.081-0.075	D03
52.71	10.00	9.215+0.151-0.133	D03
68.33	10.00	9.919+0.291-0.229	D03
83.95	10.00	10.177+0.356-0.268	D03
99.56	10.00	> 10.0	D03
115.63	10.00	> 10.1	D03
131.25	10.00	> 10.1	D03

RAPTOR-S							

135.46	5.00	10.713+0.011	
144.56	5.00	10.775+0.011	
153.66	5.00	10.854+0.011	
162.76	5.00	10.874+0.011	
171.86	5.00	10.923+0.011	
180.86	5.00	10.949+0.011	
189.96	5.00	10.999+0.011	
199.06	5.00	11.021+0.012	
208.16	5.00	11.048+0.012	
221.96	10.00	11.135+0.011	
234.76	10.00	11.177+0.011	
247.96	10.00	11.245+0.011	
261.17	10.00	11.290+0.011	
273.96	10.00	11.341+0.011	

287.16	10.00	11.382+-0.011
299.97	10.00	11.445+-0.011
313.16	10.00	11.486+-0.011
325.97	10.00	11.558+-0.011
338.76	10.00	11.596+-0.011
351.96	10.00	11.631+-0.012
364.76	10.00	11.664+-0.012
377.56	10.00	11.721+-0.012
390.36	10.00	11.773+-0.012
403.16	10.00	11.814+-0.012
416.36	10.00	11.853+-0.012
429.16	10.00	11.877+-0.012
441.96	10.00	11.917+-0.012
454.76	10.00	11.949+-0.012
467.56	10.00	12.010+-0.012
492.56	30.00	12.049+-0.011
534.86	30.00	12.147+-0.011
577.16	30.00	12.222+-0.011
619.56	30.00	12.294+-0.011
661.96	30.00	12.380+-0.011
704.26	30.00	12.469+-0.011
746.66	30.00	12.526+-0.012
788.96	30.00	12.579+-0.012
831.76	30.00	12.663+-0.012
874.16	30.00	12.725+-0.012
916.46	30.00	12.789+-0.012
958.46	30.00	12.858+-0.012
1001.16	30.00	12.916+-0.012
1043.16	30.00	12.993+-0.012
1085.56	30.00	13.034+-0.012
1128.37	30.00	13.097+-0.012
1170.86	30.00	13.154+-0.012
1213.16	30.00	13.178+-0.013
1255.96	30.00	13.237+-0.013
1297.96	30.00	13.279+-0.013
1339.96	30.00	13.318+-0.013
1382.76	30.00	13.357+-0.013
1425.16	30.00	13.379+-0.013
1467.46	30.00	13.446+-0.013
1509.76	30.00	13.471+-0.013
1552.36	30.00	13.493+-0.013
1594.76	30.00	13.535+-0.013
1637.26	30.00	13.560+-0.013
1679.66	30.00	13.593+-0.013
1721.96	30.00	13.626+-0.014
1764.66	30.00	13.652+-0.014
1806.96	30.00	13.666+-0.014
1849.37	30.00	13.680+-0.014
1891.76	30.00	13.730+-0.014
1933.86	30.00	13.740+-0.014
1975.86	30.00	13.756+-0.014

2018.26	30.00	13.790+-0.014
2060.36	30.00	13.813+-0.014
2102.36	30.00	13.833+-0.014
2144.66	30.00	13.851+-0.014
2187.36	30.00	13.895+-0.015
2229.36	30.00	13.904+-0.015
2271.86	30.00	13.934+-0.015
2314.26	30.00	13.954+-0.015
2356.66	30.00	13.966+-0.015
2398.66	30.00	13.984+-0.015
2440.76	30.00	14.004+-0.015
2483.16	30.00	14.042+-0.015
2525.56	30.00	14.052+-0.015
2568.26	30.00	14.071+-0.015
2611.07	30.00	14.092+-0.015
2653.36	30.00	14.097+-0.015
2696.16	30.00	14.127+-0.015
2738.16	30.00	14.152+-0.016
2780.36	30.00	14.151+-0.016
2822.36	30.00	14.178+-0.016
2864.76	30.00	14.188+-0.016
2907.16	30.00	14.185+-0.016
2949.46	30.00	14.206+-0.016
2991.46	30.00	14.221+-0.016
3033.86	30.00	14.241+-0.016
3075.86	30.00	14.276+-0.016
3118.26	30.00	14.287+-0.016
3160.56	30.00	14.296+-0.017
3203.36	30.00	14.332+-0.017
3245.46	30.00	14.322+-0.017
3288.26	30.00	14.324+-0.017
3330.46	30.00	14.342+-0.017
3372.76	30.00	14.336+-0.017
3415.56	30.00	14.391+-0.017
3457.56	30.00	14.389+-0.017
3499.97	30.00	14.383+-0.017
3542.66	30.00	14.380+-0.018
3585.06	30.00	14.408+-0.018
3627.66	30.00	14.431+-0.018
3669.66	30.00	14.432+-0.018
3711.96	30.00	14.450+-0.018
3754.26	30.00	14.455+-0.018
3796.96	30.00	14.467+-0.018
3839.46	30.00	14.489+-0.018
3881.76	30.00	14.510+-0.018
3924.46	30.00	14.504+-0.018
3966.87	30.00	14.526+-0.019
4008.86	30.00	14.514+-0.018
4051.56	30.00	14.525+-0.018
4093.96	30.00	14.578+-0.019
4136.36	30.00	14.558+-0.019

4178.56	30.00	14.587+-0.019
4221.26	30.00	14.587+-0.019
4263.46	30.00	14.583+-0.019
4305.86	30.00	14.592+-0.019
4348.66	30.00	14.574+-0.019
4391.06	30.00	14.596+-0.019
4433.37	30.00	14.604+-0.019
4476.06	30.00	14.606+-0.019
4518.16	30.00	14.605+-0.020
4560.46	30.00	14.651+-0.020
4602.66	30.00	14.634+-0.020
4645.06	30.00	14.664+-0.020
4687.56	30.00	14.649+-0.020
4730.26	30.00	14.674+-0.020
4772.46	30.00	14.655+-0.020
4814.86	30.00	14.679+-0.020
4857.26	30.00	14.670+-0.021
4899.36	30.00	14.717+-0.021
4941.66	30.00	14.723+-0.020
4983.66	30.00	14.719+-0.021
5026.06	30.00	14.772+-0.022

RAPTOR-T

138.20	5.00	11.098+-0.019	10.822+-0.009	10.665+-0.013	10.498+-0.025
149.62	5.00	11.179+-0.019	10.894+-0.009	10.716+-0.013	10.565+-0.025
161.03	5.00	11.231+-0.019	10.934+-0.010	10.800+-0.013	10.675+-0.025
172.44	5.00	11.278+-0.019	10.974+-0.010	10.859+-0.013	10.716+-0.025
183.85	5.00	11.263+-0.019	11.008+-0.010	10.899+-0.013	10.780+-0.025
195.66	5.00	11.337+-0.020	11.080+-0.010	10.955+-0.013	10.786+-0.025
207.58	5.00	11.337+-0.020	11.108+-0.010	10.971+-0.013	10.875+-0.025
218.99	5.00	11.447+-0.020	11.159+-0.010	11.037+-0.013	10.902+-0.025
230.81	5.00	11.451+-0.020	11.173+-0.010	11.088+-0.013	10.893+-0.026
247.24	10.00	11.582+-0.019	11.289+-0.009	11.178+-0.013	10.986+-0.025
260.61	10.00	11.585+-0.019	11.319+-0.009	11.178+-0.013	11.027+-0.024
273.60	10.00	11.631+-0.019	11.359+-0.009	11.237+-0.013	11.127+-0.024
286.97	10.00	11.687+-0.019	11.406+-0.009	11.290+-0.013	11.139+-0.024
299.96	10.00	11.745+-0.019	11.467+-0.009	11.337+-0.013	11.200+-0.025
312.98	10.00	11.822+-0.019	11.520+-0.009	11.414+-0.013	11.238+-0.025
325.94	10.00	11.867+-0.019	11.555+-0.010	11.456+-0.013	11.300+-0.025
338.93	10.00	11.909+-0.019	11.606+-0.010	11.497+-0.013	11.366+-0.025
352.30	10.00	11.933+-0.019	11.672+-0.010	11.545+-0.013	11.389+-0.025
365.69	10.00	12.032+-0.020	11.709+-0.010	11.604+-0.013	11.436+-0.025
379.07	10.00	12.064+-0.020	11.770+-0.010	11.620+-0.013	11.469+-0.025
392.46	10.00	12.088+-0.020	11.801+-0.010	11.667+-0.013	11.527+-0.025
405.83	10.00	12.135+-0.020	11.842+-0.010	11.716+-0.013	11.571+-0.025
419.01	10.00	12.171+-0.020	11.878+-0.010	11.735+-0.013	11.636+-0.026
431.99	10.00	12.212+-0.020	11.925+-0.010	11.803+-0.013	11.606+-0.026
445.17	10.00	12.261+-0.020	11.957+-0.010	11.825+-0.013	11.671+-0.026
458.55	10.00	12.283+-0.020	11.974+-0.010	11.870+-0.014	11.729+-0.026
472.13	10.00	12.302+-0.020	12.028+-0.010	11.883+-0.014	11.718+-0.026

485.10	10.00	12.342+-0.020	12.058+-0.010	11.925+-0.014	11.757+-0.026
498.08	10.00	12.399+-0.020	12.101+-0.010	11.960+-0.014	11.779+-0.026
525.15	30.00	12.458+-0.019	12.153+-0.009	12.016+-0.013	11.864+-0.024
567.05	30.00	12.526+-0.019	12.257+-0.009	12.119+-0.013
608.32	30.00	12.610+-0.019	12.324+-0.009	12.188+-0.013
650.06	30.00	12.696+-0.019	12.407+-0.009	12.270+-0.013
691.73	30.00	12.763+-0.019	12.464+-0.009	12.335+-0.013
733.47	30.00	12.822+-0.019	12.528+-0.009	12.396+-0.013
774.78	30.00	12.888+-0.020	12.589+-0.010	12.481+-0.013
816.03	30.00	12.949+-0.020	12.670+-0.010	12.529+-0.013
857.68	30.00	13.021+-0.020	12.749+-0.010	12.579+-0.013
898.98	30.00	13.073+-0.020	12.800+-0.010	12.659+-0.013
940.65	30.00	13.153+-0.020	12.877+-0.010	12.713+-0.013
982.35	30.00	13.241+-0.020	12.941+-0.010	12.780+-0.014
1023.70	30.00	13.280+-0.020	12.996+-0.010	12.873+-0.014
1065.60	30.00	13.365+-0.020	13.063+-0.010	12.901+-0.014
1107.42	30.00	13.391+-0.020	13.105+-0.010	12.978+-0.014
1148.82	30.00	13.453+-0.020	13.141+-0.010	13.029+-0.014
1190.49	30.00	13.486+-0.020	13.199+-0.010	13.052+-0.014
1231.82	30.00	13.539+-0.020	13.234+-0.011	13.117+-0.014
1273.43	30.00	13.588+-0.021	13.289+-0.011	13.162+-0.015
1315.24	30.00	13.632+-0.021	13.330+-0.011	13.213+-0.015
1357.05	30.00	13.651+-0.021	13.371+-0.011	13.253+-0.015
1399.05	30.00	13.683+-0.021	13.418+-0.011	13.271+-0.016
1440.96	30.00	13.739+-0.021	13.462+-0.011	13.324+-0.016
1483.47	30.00	13.766+-0.021	13.493+-0.011	13.351+-0.016
1525.18	30.00	13.787+-0.021	13.518+-0.011	13.377+-0.016
1566.88	30.00	13.830+-0.021	13.543+-0.011	13.428+-0.016
1608.69	30.00	13.856+-0.021	13.593+-0.011	13.438+-0.016
1649.99	30.00	13.886+-0.021	13.596+-0.011	13.470+-0.016
1691.60	30.00	13.934+-0.021	13.639+-0.012	13.514+-0.017
1733.71	30.00	13.961+-0.021	13.675+-0.012	13.534+-0.017
1775.52	30.00	13.984+-0.021	13.675+-0.012	13.530+-0.017
1816.78	30.00	13.989+-0.022	13.711+-0.012	13.580+-0.018
1858.48	30.00	14.007+-0.022	13.731+-0.012	13.603+-0.018
1900.33	30.00	14.039+-0.022	13.756+-0.012	13.649+-0.018
1941.84	30.00	14.064+-0.022	13.781+-0.013	13.655+-0.018
1983.54	30.00	14.088+-0.022	13.825+-0.013	13.704+-0.018
2024.94	30.00	14.108+-0.022	13.845+-0.013	13.716+-0.018
2066.60	30.00	14.145+-0.022	13.863+-0.013	13.737+-0.018
2108.86	30.00	14.149+-0.022	13.872+-0.013	13.774+-0.018
2150.16	30.00	14.180+-0.022	13.914+-0.013	13.754+-0.018
2191.82	30.00	14.185+-0.022	13.921+-0.013	13.776+-0.018
2233.87	30.00	14.213+-0.023	13.947+-0.013	13.809+-0.019
2275.28	30.00	14.243+-0.023	13.972+-0.013	13.867+-0.020
2316.68	30.00	14.262+-0.023	14.005+-0.013	13.872+-0.020
2358.59	30.00	14.270+-0.023	13.998+-0.013	13.885+-0.020
2400.39	30.00	14.317+-0.024	14.039+-0.013	13.901+-0.020
2442.10	30.00	14.337+-0.024	14.048+-0.013	13.894+-0.020
2483.80	30.00	14.327+-0.024	14.051+-0.014	13.941+-0.021
2525.61	30.00	14.369+-0.024	14.091+-0.014	13.934+-0.021

2567.27	30.00	14.388+-0.024	14.101+-0.014	13.986+-0.021
2608.98	30.00	14.366+-0.024	14.145+-0.014	14.026+-0.022
2650.73	30.00	14.422+-0.024	14.119+-0.014	14.029+-0.022
2692.53	30.00	14.428+-0.025	14.173+-0.015	14.048+-0.022
2734.20	30.00	14.457+-0.025	14.175+-0.015	14.030+-0.022
2775.54	30.00	14.445+-0.025	14.208+-0.015	14.072+-0.022
2817.20	30.00	14.494+-0.025	14.197+-0.015	14.094+-0.022
2858.51	30.00	14.498+-0.025	14.216+-0.015	14.095+-0.022
2900.66	30.00	14.524+-0.025	14.258+-0.016	14.061+-0.022
2942.67	30.00	14.530+-0.025	14.268+-0.016	14.111+-0.022
2984.57	30.00	14.564+-0.026	14.253+-0.016	14.149+-0.023
3025.88	30.00	14.567+-0.026	14.268+-0.016	14.136+-0.023
3067.48	30.00	14.552+-0.026	14.289+-0.016	14.175+-0.024
3109.39	30.00	14.590+-0.026	14.289+-0.016	14.188+-0.024
3150.99	30.00	14.586+-0.026	14.320+-0.017	14.319+-0.025
3192.50	30.00	14.586+-0.027	14.366+-0.017	14.194+-0.024
3234.51	30.00	14.606+-0.026	14.371+-0.017	14.248+-0.025
3276.52	30.00	14.605+-0.027	14.347+-0.017	14.242+-0.025
3317.92	30.00	14.660+-0.027	14.377+-0.017	14.228+-0.025
3359.58	30.00	14.660+-0.027	14.394+-0.017	14.312+-0.026
3401.29	30.00	14.673+-0.028	14.397+-0.017	14.274+-0.025
3443.00	30.00	14.651+-0.027	14.443+-0.017	14.431+-0.027
3484.70	30.00	14.701+-0.028	14.438+-0.017	14.309+-0.025
3526.55	30.00	14.705+-0.028	14.417+-0.017	14.311+-0.026
3568.56	30.00	14.748+-0.028	14.458+-0.017	14.455+-0.028
3610.37	30.00	14.704+-0.028	14.466+-0.017	14.361+-0.027
3651.66	30.00	14.740+-0.028	14.490+-0.018
3693.17	30.00	14.826+-0.030	14.509+-0.018
3735.08	30.00	14.773+-0.028	14.513+-0.018
3776.79	30.00	14.753+-0.028	14.498+-0.018
3818.70	30.00	14.751+-0.028	14.503+-0.018
3860.10	30.00	14.759+-0.029	14.523+-0.018
3901.50	30.00	14.789+-0.029	14.543+-0.019
3943.61	30.00	14.759+-0.028	14.531+-0.018
3985.32	30.00	14.819+-0.030	14.562+-0.019
4027.12	30.00	14.823+-0.030	14.546+-0.019
4068.78	30.00	14.843+-0.031	14.568+-0.019
4110.33	30.00	14.880+-0.031	14.595+-0.020
4152.34	30.00	14.843+-0.031	14.576+-0.019
4194.05	30.00	14.842+-0.031	14.590+-0.020
4237.97	30.00	14.885+-0.031	14.615+-0.020
4279.78	30.00	14.873+-0.031	14.643+-0.020
4321.08	30.00	14.868+-0.031	14.651+-0.020
4362.49	30.00	14.913+-0.031	14.619+-0.020
4404.39	30.00	14.950+-0.032	14.635+-0.020
4446.10	30.00	14.893+-0.031	14.656+-0.021
4487.47	30.00	14.950+-0.032	14.694+-0.021
4529.11	30.00	14.913+-0.031	14.705+-0.021
4571.01	30.00	14.950+-0.033	14.691+-0.021
4613.02	30.00	14.971+-0.033	14.691+-0.021
4654.32	30.00	14.990+-0.034	14.694+-0.022

4695.73	30.00	14.996+-0.034	14.732+-0.022
4737.33	30.00	14.978+-0.033	14.710+-0.022
4778.94	30.00	14.954+-0.033	14.707+-0.022
4820.34	30.00	14.965+-0.033	14.754+-0.023
4862.15	30.00	14.978+-0.033	14.722+-0.022
4903.65	30.00	14.968+-0.034	14.743+-0.023
4944.95	30.00	15.035+-0.035	14.752+-0.023
4986.76	30.00	15.027+-0.035	14.795+-0.023
5028.77	30.00	15.013+-0.035	14.785+-0.024
5070.48	30.00	15.026+-0.035	14.808+-0.024
5112.19	30.00	15.019+-0.035	14.810+-0.024
5153.79	30.00	15.064+-0.036	14.824+-0.025
5195.70	30.00	15.072+-0.036	14.802+-0.024
5236.99	30.00	15.071+-0.036	14.818+-0.024
5278.26	30.00	15.063+-0.036	14.867+-0.025
5320.31	30.00	15.051+-0.036	14.833+-0.024
5362.02	30.00	15.090+-0.037	14.818+-0.025
5403.82	30.00	15.040+-0.036	14.877+-0.026
5445.63	30.00	15.123+-0.038	14.886+-0.026
5487.53	30.00	15.110+-0.037	14.889+-0.026
5528.94	30.00	15.165+-0.039	14.868+-0.026
5570.24	30.00	15.099+-0.038	14.866+-0.026
5611.90	30.00	15.074+-0.036	14.873+-0.026
5653.65	30.00	15.159+-0.039	14.888+-0.026
5695.76	30.00	15.186+-0.039	14.901+-0.026
5737.22	30.00	15.189+-0.040	14.905+-0.027
5779.17	30.00	15.185+-0.040	14.927+-0.027
5820.98	30.00	15.192+-0.040	14.927+-0.028
5862.68	30.00	15.163+-0.040	14.910+-0.027
5904.39	30.00	15.194+-0.041	14.891+-0.027
5946.10	30.00	15.194+-0.041	14.943+-0.028
5987.80	30.00	15.183+-0.041	14.975+-0.028
6029.11	30.00	15.177+-0.041	14.909+-0.028
6070.91	30.00	15.169+-0.041	14.970+-0.028
6112.58	30.00	15.125+-0.039	14.988+-0.028
6154.28	30.00	15.231+-0.042	14.986+-0.028
6195.99	30.00	15.204+-0.042	14.983+-0.028
6237.43	30.00	15.212+-0.042	14.977+-0.028
6279.34	30.00	15.238+-0.044	14.965+-0.028
6320.74	30.00	15.256+-0.044	14.986+-0.029
6362.24	30.00	15.257+-0.045	15.036+-0.030
6403.95	30.00	15.208+-0.042	15.007+-0.029
6445.76	30.00	15.258+-0.044	14.984+-0.029
6487.46	30.00	15.215+-0.042	15.035+-0.030
6529.27	30.00	15.367+-0.048	15.002+-0.029
6570.97	30.00	15.256+-0.044	15.044+-0.031
6612.28	30.00	15.329+-0.049	15.057+-0.032
6653.58	30.00	15.242+-0.045	15.080+-0.032
6695.29	30.00	15.245+-0.044	15.059+-0.031
6736.54	30.00	15.322+-0.047	15.057+-0.032
6777.89	30.00	15.295+-0.047	15.037+-0.032

6819.29	30.00	15.340+-0.048	15.053+-0.032
6861.10	30.00	15.276+-0.046	15.034+-0.031
6903.00	30.00	15.294+-0.047	15.080+-0.033
6944.82	30.00	15.294+-0.047	15.024+-0.031
6986.47	30.00	15.224+-0.045	15.117+-0.033
7028.33	30.00	15.355+-0.049	15.148+-0.035
7069.73	30.00	15.372+-0.050	15.066+-0.033
7111.44	30.00	15.348+-0.051	15.132+-0.034
7152.94	30.00	15.340+-0.051	15.100+-0.034
7194.75	30.00	15.283+-0.048	15.076+-0.033
7236.45	30.00	15.378+-0.051	15.130+-0.035
7278.16	30.00	15.354+-0.051	15.124+-0.035
7319.96	30.00	15.425+-0.053	15.161+-0.036
7361.67	30.00	15.396+-0.053	15.182+-0.037
7403.37	30.00	15.333+-0.051	15.151+-0.036
7445.39	30.00	15.458+-0.055	15.185+-0.037
7487.39	30.00	15.428+-0.055	15.106+-0.035
7529.30	30.00	15.337+-0.052	15.166+-0.036
7570.90	30.00	15.475+-0.058	15.173+-0.037

References and Notes

1. S. E. Woosley, J. S. Bloom, The supernova gamma-ray burst connection. *Annu. Rev. Astron. Astrophys.* **44**, 507–556 (2006). [doi:10.1146/annurev.astro.43.072103.150558](https://doi.org/10.1146/annurev.astro.43.072103.150558)
2. B. D. Metzger, D. Giannios, T. A. Thompson, N. Bucciantini, E. Quataert, The protomagnetar model for gamma-ray bursts. *Mon. Not. R. Astron. Soc.* **413**, 2031–2056 (2011). [doi:10.1111/j.1365-2966.2011.18280.x](https://doi.org/10.1111/j.1365-2966.2011.18280.x)
3. M. R. Metzger, S. G. Djorgovski, S. R. Kulkarni, C. C. Steidel, K. L. Adelberger, D. A. Frail, E. Costa, F. Frontera, Spectral constraints on the redshift of the optical counterpart to the gamma-ray burst of 8 May 1997. *Nature* **387**, 878 (1997). [doi:10.1038/43132](https://doi.org/10.1038/43132)
4. A. Maselli, A. Melandri, L. Nava, C. G. Mundell, N. Kawai, S. Campana, S. Covino, J. R. Cummings, G. Cusumano, P. A. Evans, G. Ghirlanda, G. Ghisellini, C. Guidorzi, S. Kobayashi, P. Kuin, V. La Parola, V. Mangano, S. Oates, T. Sakamoto, M. Serino, F. Virgili, B.-B. Zhang, S. Barthelmy, A. Beardmore, M. G. Bernardini, D. Bersier, D. Burrows, G. Calderone, M. Capalbi, J. Chiang, P. D’Avanzo, V. D’Elia, M. De Pasquale, D. Fugazza, N. Gehrels, A. Gomboc, R. Harrison, H. Hanayama, J. Japelj, J. Kennea, D. Kopac, C. Kouveliotou, D. Kuroda, A. Levan, D. Malesani, F. Marshall, J. Nousek, P. O’Brien, J. P. Osborne, C. Pagani, K. L. Page, M. Page, M. Perri, T. Pritchard, P. Romano, Y. Saito, B. Sbarufatti, R. Salvaterra, I. Steele, N. Tanvir, G. Vianello, B. Weigand, K. Wiersema, Y. Yatsu, T. Yoshii, G. Tagliaferri, GRB 130427A: A nearby ordinary monster. *Science* 10.1126/science.1242279 (2013).
5. M. Ackermann, M. Ajello, K. Asano, W. B. Atwood, M. Axelsson, L. Baldini, J. Ballet, G. Barbiellini, M. G. Baring, D. Bastieri, K. Bechtol, R. Bellazzini, E. Bissaldi, E. Bonamente, J. Bregeon, M. Brigida, P. Bruel, R. Buehler, J. Michael Burgess, S. Buson, G. A. Caliendo, R. A. Cameron, P. A. Caraveo, C. Cecchi, V. Chaplin, E. Charles, A. Chekhtman, C. C. Cheung, J. Chiang, G. Chiaro, S. Ciprini, R. Claus, W. Cleveland, J. Cohen-Tanugi, A. Collazzi, L. R. Cominsky, V. Connaughton, J. Conrad, S. Cutini, F. D’Ammando, A. de Angelis, M. DeKlotz, F. de Palma, C. D. Dermer, R. Desiante, A. Diekmann, L. Di Venere, P. S. Drell, A. Drlica-Wagner, C. Favuzzi, S. J. Fegan, E. C. Ferrara, J. Finke, G. Fitzpatrick, W. B. Focke, A. Franckowiak, Y. Fukazawa, S. Funk, P. Fusco, F. Gargano, N. Gehrels, S. Germani, M. Gibby, N. Giglietto, M. Giles, F. Giordano, M. Giroletti, G. Godfrey, J. Granot, I. A. Grenier, J. E. Grove, D. Gruber, S. Guiriec, D. Hadasch, Y. Hanabata, A. K. Harding, M. Hayashida, E. Hays, D. Horan, R. E. Hughes, Y. Inoue, T. Jogler, G. Jóhannesson, W. N. Johnson, T. Kawano, J. Knödseder, D. Kocevski, M. Kuss, J. Lande, S. Larsson, L. Latronico, F. Longo, F. Loparco, M. N. Lovellette, P. Lubrano, M. Mayer, M. N. Mazziotta, J. E. McEnery, P. F. Michelson, T. Mizuno, A. A. Moiseev, M. E. Monzani, E. Moretti, A. Morselli, I. V. Moskalenko, S. Murgia, R. Nemmen, E. Nuss, M. Ohno, T. Ohsugi, A. Okumura, N. Omodei, M. Orienti, D. Paneque, V. Pelassa, J. S. Perkins, M. Pesce-Rollins, V. Petrosian, F. Piron, G. Pivato, T. A. Porter, J. L. Racusin, S. Rainò, R. Rando, M. Razzano, S. Razzaque, A. Reimer, O. Reimer, S. Ritz, M. Roth, F. Ryde, A. Sartori, P. M. Saz Parkinson, J. D. Scargle, A. Schulz, C. Sgrò, E. J. Siskind, E. Sonbas, G. Spandre, P. Spinelli, H. Tajima, H. Takahashi, J. G. Thayer, J. B. Thayer, D. J. Thompson, L. Tibaldo, M. Tinivella, D. F. Torres, G. Tosti, E. Troja, T. L. Usher, J. Vandenbroucke, V. Vasileiou, G. Vianello, V. Vitale, B. L. Winer, K. S. Wood, R. Yamazaki, G. Younes, H.-

- F. Yu, S. J. Zhu, P. N. Bhat, M. S. Briggs, D. Byrne, S. Foley, A. Goldstein, P. Jenke, R. M. Kippen, C. Kouveliotou, S. McBreen, C. Meegan, W. S. Paciesas, R. Preece, A. Rau, D. Tierney, A. J. van der Horst, A. von Kienlin, C. Wilson-Hodge, S. Xiong, G. Cusumano, V. La Parola, J. R. Cummings, Fermi-LAT observations of the gamma-ray burst GRB 130427A. *Science* 10.1126/science.1242353 (2013).
6. S. Golenetskii *et al.*, “Konus-wind observation of GRB 130427A,” GCN Circular #14486 (2013).
 7. A. J. Levan *et al.*, “GRB 130427A: Gemini-north redshift,” GCN circular #14455 (2013).
 8. D. Xu, A. de Ugarte Postigo, G. Leloudas, T. Krühler, Z. Cano, J. Hjorth, D. Malesani, J. P. U. Fynbo, C. C. Thöne, R. Sánchez-Ramírez, S. Schulze, P. Jakobsson, L. Kaper, J. Sollerman, D. J. Watson, A. Cabrera-Lavers, C. Cao, S. Covino, H. Flores, S. Geier, J. Gorosabel, S. M. Hu, B. Milvang-Jensen, M. Sparre, L. P. Xin, T. M. Zhang, W. K. Zheng, Y. C. Zou, Discovery of the broad-lined Type Ic SN 2013cq associated with the very energetic GRB 130427A. *Astrophys. J.* **776**, 98 (2013). [doi:10.1088/0004-637X/776/2/98](https://doi.org/10.1088/0004-637X/776/2/98)
 9. J. Wren, W. T. Vestrand, P. Wozniak, H. Davis, A portable observatory for persistent monitoring of the night sky. *Proc. SPIE* **7737**, 773723 (2010). [doi:10.1117/12.859039](https://doi.org/10.1117/12.859039)
 10. W. T. Vestrand, P. R. Wozniak, J. A. Wren, E. E. Fenimore, T. Sakamoto, R. R. White, D. Casperson, H. Davis, S. Evans, M. Galassi, K. E. McGowan, J. A. Schier, J. W. Asa, S. D. Barthelmy, J. R. Cummings, N. Gehrels, D. Hullinger, H. A. Krimm, C. B. Markwardt, K. McLean, D. Palmer, A. Parsons, J. Tueller, A link between prompt optical and prompt gamma-ray emission in gamma-ray bursts. *Nature* **435**, 178–180 (2005). [Medline doi:10.1038/nature03515](https://doi.org/10.1038/nature03515)
 11. W. T. Vestrand, J. A. Wren, P. R. Wozniak, R. Aptekar, S. Golenetskii, V. Pal’shin, T. Sakamoto, R. R. White, S. Evans, D. Casperson, E. Fenimore, Energy input and response from prompt and early optical afterglow emission in gamma-ray bursts. *Nature* **442**, 172–175 (2006). [Medline doi:10.1038/nature04913](https://doi.org/10.1038/nature04913)
 12. J. L. Racusin, S. V. Karpov, M. Sokolowski, J. Granot, X. F. Wu, V. Pal’shin, S. Covino, A. J. van der Horst, S. R. Oates, P. Schady, R. J. Smith, J. Cummings, R. L. Starling, L. W. Piotrowski, B. Zhang, P. A. Evans, S. T. Holland, K. Malek, M. T. Page, L. Vetere, R. Margutti, C. Guidorzi, A. P. Kamble, P. A. Curran, A. Beardmore, C. Kouveliotou, L. Mankiewicz, A. Melandri, P. T. O’Brien, K. L. Page, T. Piran, N. R. Tanvir, G. Wrochna, R. L. Aptekar, S. Barthelmy, C. Bartolini, G. M. Beskin, S. Bondar, M. Bremer, S. Campana, A. Castro-Tirado, A. Cucchiara, M. Cwiok, P. D’Avanzo, V. D’Elia, M. D. Valle, A. de Ugarte Postigo, W. Dominik, A. Falcone, F. Fiore, D. B. Fox, D. D. Frederiks, A. S. Fruchter, D. Fugazza, M. A. Garrett, N. Gehrels, S. Golenetskii, A. Gomboc, J. Gorosabel, G. Greco, A. Guarnieri, S. Immler, M. Jelinek, G. Kasprovicz, V. La Parola, A. J. Levan, V. Mangano, E. P. Mazets, E. Molinari, A. Moretti, K. Nawrocki, P. P. Oleynik, J. P. Osborne, C. Pagani, S. B. Pandey, Z. Paragi, M. Perri, A. Piccioni, E. Ramirez-Ruiz, P. W. Roming, I. A. Steele, R. G. Strom, V. Testa, G. Tosti, M. V. Ulanov, K. Wiersema, R. A. Wijers, J. M. Winters, A. F. Zarnecki, F. Zerbi, P. Mészáros, G. Chincarini, D. N. Burrows, Broadband observations of the naked-eye gamma-ray burst GRB 080319B. *Nature* **455**, 183–188 (2008). [Medline doi:10.1038/nature07270](https://doi.org/10.1038/nature07270)

13. C. Akerlof, R. Balsano, S. Barthelmy, J. Bloch, P. Butterworth, D. Casperson, T. Cline, S. Fletcher, F. Frontera, G. Gisler, J. Heise, J. Hills, R. Kehoe, B. Lee, S. Marshall, T. McKay, R. Miller, L. Piro, W. Priedhorsky, J. Szymanski, J. Wren, Observations of contemporaneous optical radiation from a gamma-ray burst. *Nature* **398**, 400–402 (1999). [doi:10.1038/18837](https://doi.org/10.1038/18837)
14. E. Rykoff, F. Aharonian, C. W. Akerlof, M. C. B. Ashley, S. D. Barthelmy, H. A. Flewelling, N. Gehrels, E. Göğüş, T. Güver, Ü. Kiziloğlu, H. A. Krimm, T. A. McKay, M. Özel, A. Phillips, R. M. Quimby, G. Rowell, W. Rujopakarn, B. E. Schaefer, D. A. Smith, W. T. Vestrand, J. C. Wheeler, J. Wren, F. Yuan, S. A. Yost, Looking into the Fireball: ROTSE-III and *Swift* observations of early gamma-ray burst afterglows. *Astrophys. J.* **702**, 489–505 (2009). [doi:10.1088/0004-637X/702/1/489](https://doi.org/10.1088/0004-637X/702/1/489)
15. P. Meszaros, M. Rees, Relativistic fireballs and their impact on external matter—Models for cosmological gamma-ray bursts. *Astrophys. J.* **405**, 278 (1993). [doi:10.1086/172360](https://doi.org/10.1086/172360)
16. P. Meszaros, M. Rees, Optical and long wavelength afterglow from gamma-ray bursts. *Astrophys. J.* **476**, 232–237 (1997). [doi:10.1086/303625](https://doi.org/10.1086/303625)
17. R. Sari, T. Piran, Predictions for the very early afterglow and the optical flash. *Astrophys. J.* **520**, 641–649 (1999). [doi:10.1086/307508](https://doi.org/10.1086/307508)
18. T. Laskar, E. Berger, B. A. Zauderer, R. Margutti, A. M. Soderberg, S. Chakraborti, R. Lunnan, R. Chornock, P. Chandra, A. Ray, A reverse shock in GRB 130427A. available at <http://arxiv.org/abs/1305.2453> (2013).
19. Z. L. Uhm, B. Zhang, R. Hascoët, F. Daigne, R. Mochkovitch, I. H. Park, Dynamics and afterglow light curves of gamma-ray burst blast waves with a long-lived reverse shock. *Astrophys. J.* **761**, 147 (2012). [doi:10.1088/0004-637X/761/2/147](https://doi.org/10.1088/0004-637X/761/2/147)
20. P. Wozniak, W. T. Vestrand, A. D. Panaitescu, J. A. Wren, H. R. Davis, R. R. White, Gamma-Ray burst at the extreme: The naked-eye burst GRB 080319B. *Astrophys. J.* **691**, 495–502 (2009). [doi:10.1088/0004-637X/691/1/495](https://doi.org/10.1088/0004-637X/691/1/495)
21. J. S. Bloom, D. A. Perley, W. Li, N. R. Butler, A. A. Miller, D. Kocevski, D. A. Kann, R. J. Foley, H.-W. Chen, A. V. Filippenko, D. L. Starr, B. Macomber, J. X. Prochaska, R. Chornock, D. Poznanski, S. Klose, M. F. Skrutskie, S. Lopez, P. Hall, K. Glazebrook, C. H. Blake, Observations of the naked-eye GRB 080319B: implications of nature's brightest explosion. *Astrophys. J.* **691**, 723–737 (2009). [doi:10.1088/0004-637X/691/1/723](https://doi.org/10.1088/0004-637X/691/1/723)
22. T. W. Giblin, J. van Paradijs, C. Kouveliotou, V. Connaughton, R. A. M. J. Wijers, M. S. Briggs, R. D. Preece, G. J. Fishman, Evidence for an early high-energy afterglow observed with BATSE from GRB 980923. *Astrophys. J.* **524**, L47–L50 (1999). [doi:10.1086/312285](https://doi.org/10.1086/312285)
23. A. Panaitescu, W. T. Vestrand, Taxonomy of gamma-ray burst optical light curves: Identification of a salient class of early afterglows. *Mon. Not. R. Astron. Soc.* **387**, 497–504 (2008). [doi:10.1111/j.1365-2966.2008.13231.x](https://doi.org/10.1111/j.1365-2966.2008.13231.x)
24. The RAPTOR optical measurements are available in table S1 of the supplementary materials on Science Online.

25. R. Blandford, C. McKee, Fluid dynamics of relativistic blast waves. *Phys. Fluids* **19**, 1130 (1976). [doi:10.1063/1.861619](https://doi.org/10.1063/1.861619)
26. E. Bertin, S. Arnouts, SExtractor: Software for source extraction. *Astron. Astrophys. Suppl. Ser.* **117**, 393–404 (1996). [doi:10.1051/aas:1996164](https://doi.org/10.1051/aas:1996164)
27. E. Hog *et al.*, The Tycho-2 catalogue of the 2.5 million brightest stars. *Astron. Astrophys.* **355**, L27 (2000).
28. E. Pickles, E. Depagne, All sky spectrally matched *UBVRI-ZY* and *u'g'r'i'z'* magnitudes for stars in the Tycho2 catalog. *Publ. Astron. Soc. Pac.* **122**, 1437–1464 (2010). [doi:10.1086/657947](https://doi.org/10.1086/657947)
29. C. Ahn, R. Alexandroff, C. Allende Prieto, S. F. Anderson, T. Anderton, B. H. Andrews, É. Aubourg, S. Bailey, E. Balbinot, R. Barnes, J. Bautista, T. C. Beers, A. Beifiori, A. A. Berlind, V. Bhardwaj, D. Bizyaev, C. H. Blake, M. R. Blanton, M. Blomqvist, J. J. Bochanski, A. S. Bolton, A. Borde, J. Bovy, W. N. Brandt, J. Brinkmann, P. J. Brown, J. R. Brownstein, K. Bundy, N. G. Busca, W. Carithers, A. R. Carnero, M. A. Carr, D. I. Casetti-Dinescu, Y. Chen, C. Chiappini, J. Comparat, N. Connolly, J. R. Crepp, S. Cristiani, R. A. C. Croft, A. J. Cuesta, L. N. da Costa, J. R. A. Davenport, K. S. Dawson, R. de Putter, N. De Lee, T. Delubac, S. Dhital, A. Ealet, G. L. Ebelke, E. M. Edmondson, D. J. Eisenstein, S. Escoffier, M. Esposito, M. L. Evans, X. Fan, B. Femenía Castellá, E. Fernández Alvar, L. D. Ferreira, N. Filiz Ak, H. Finley, S. W. Fleming, A. Font-Ribera, P. M. Frinchaboy, D. A. García-Hernández, A. E. G. Pérez, J. Ge, R. Génova-Santos, B. A. Gillespie, L. Girardi, J. I. González Hernández, E. K. Grebel, J. E. Gunn, H. Guo, D. Haggard, J.-C. Hamilton, D. W. Harris, S. L. Hawley, F. R. Hearty, S. Ho, D. W. Hogg, J. A. Holtzman, K. Honscheid, J. Huehnerhoff, I. I. Ivans, Ž. Ivezić, H. R. Jacobson, L. Jiang, J. Johansson, J. A. Johnson, G. Kauffmann, D. Kirkby, J. A. Kirkpatrick, M. A. Klaene, G. R. Knapp, J.-P. Kneib, J.-M. Le Goff, A. Leauthaud, K.-G. Lee, Y. S. Lee, D. C. Long, C. P. Loomis, S. Lucatello, B. Lundgren, R. H. Lupton, B. Ma, Z. Ma, N. MacDonald, C. E. Mack, S. Mahadevan, M. A. G. Maia, S. R. Majewski, M. Makler, E. Malanushenko, V. Malanushenko, A. Manchado, R. Mandelbaum, M. Manera, C. Maraston, D. Margala, S. L. Martell, C. K. McBride, I. D. McGreer, R. G. McMahon, B. Ménard, S. Meszaros, J. Miralda-Escudé, A. D. Montero-Dorta, F. Montesano, H. L. Morrison, D. Muna, J. A. Munn, H. Murayama, A. D. Myers, A. F. Neto, D. C. Nguyen, R. C. Nichol, D. L. Nidever, P. Noterdaeme, S. E. Nuza, R. L. C. Ogando, M. D. Olmstead, D. J. Oravetz, R. Owen, N. Padmanabhan, N. Palanque-Delabrouille, K. Pan, J. K. Parejko, P. Parihar, I. Pâris, P. Pattarakijwanich, J. Pepper, W. J. Percival, I. Pérez-Fournon, I. Pérez-Ràfols, P. Petitjean, J. Pforr, M. M. Pieri, M. H. Pinsonneault, G. F. Porto de Mello, F. Prada, A. M. Price-Whelan, M. J. Raddick, R. Rebolo, J. Rich, G. T. Richards, A. C. Robin, H. J. Rocha-Pinto, C. M. Rockosi, N. A. Roe, A. J. Ross, N. P. Ross, G. Rossi, J. A. Rubiño-Martín, L. Samushia, J. Sanchez Almeida, A. G. Sánchez, B. Santiago, C. Sayres, D. J. Schlegel, K. J. Schlesinger, S. J. Schmidt, D. P. Schneider, M. Schultheis, A. D. Schwobe, C. G. Scóccola, U. Seljak, E. Sheldon, Y. Shen, Y. Shu, J. Simmerer, A. E. Simmons, R. A. Skibba, M. F. Skrutskie, A. Slosar, F. Sobreira, J. S. Sobeck, K. G. Stassun, O. Steele, M. Steinmetz, M. A. Strauss, A. Streblyanska, N. Suzuki, M. E. C. Swanson, T. Tal, A. R. Thakar, D. Thomas, B. A. Thompson, J. L. Tinker, R. Tojeiro, C. A. Tremonti, M. Vargas Magaña, L. Verde, M. Viel, S. K. Vikas, N. P. Vogt, D. A. Wake, J. Wang, B. A. Weaver, D. H. Weinberg, B. J. Weiner, A. A.

- West, M. White, J. C. Wilson, J. P. Wisniewski, W. M. Wood-Vasey, B. Yanny, C. Yèche, D. G. York, O. Zamora, G. Zasowski, I. Zehavi, G.-B. Zhao, Z. Zheng, G. Zhu, J. C. Zinn, The ninth data release of the Sloan Digital Sky Survey: First spectroscopic data from the SDSS-III Baryon Oscillation Spectroscopic Survey. *Astrophys. J. Suppl. Ser.* **203**, 21 (2012). [doi:10.1088/0067-0049/203/2/21](https://doi.org/10.1088/0067-0049/203/2/21)
30. Y. C. Pei, Interstellar dust from the Milky Way to the Magellanic Clouds. *Astrophys. J.* **395**, 130 (1992). [doi:10.1086/171637](https://doi.org/10.1086/171637)
31. D. J. Schlegel, D. P. Finkbeiner, M. Davis, Maps of dust infrared emission for use in estimation of reddening and cosmic microwave background radiation foregrounds. *Astrophys. J.* **500**, 525–553 (1998). [doi:10.1086/305772](https://doi.org/10.1086/305772)